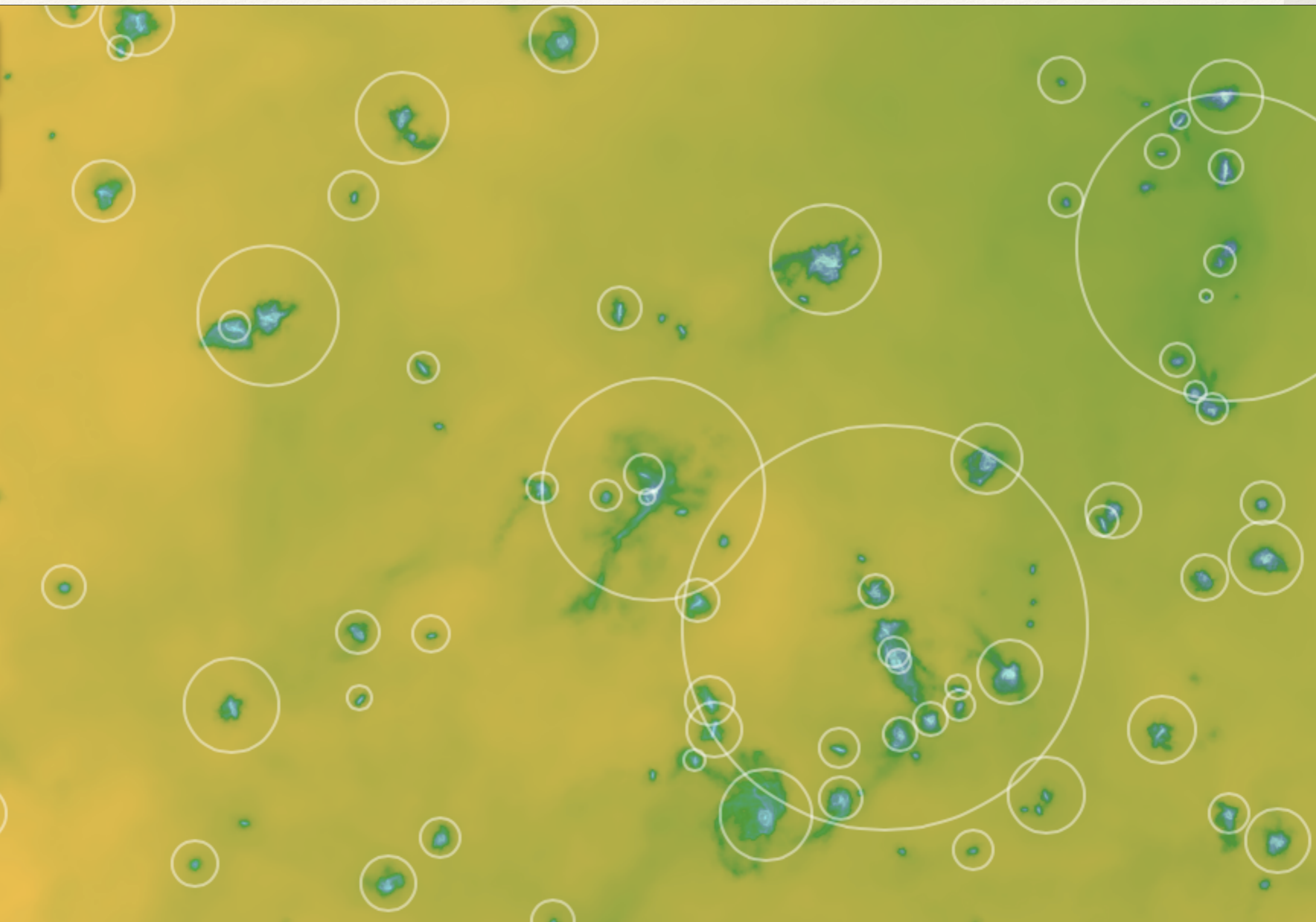


Dwarfs Structural Properties: Environmental Effects

Laura V. Sales

ITC Fellow, Harvard





Satellite Galaxies in SDSS



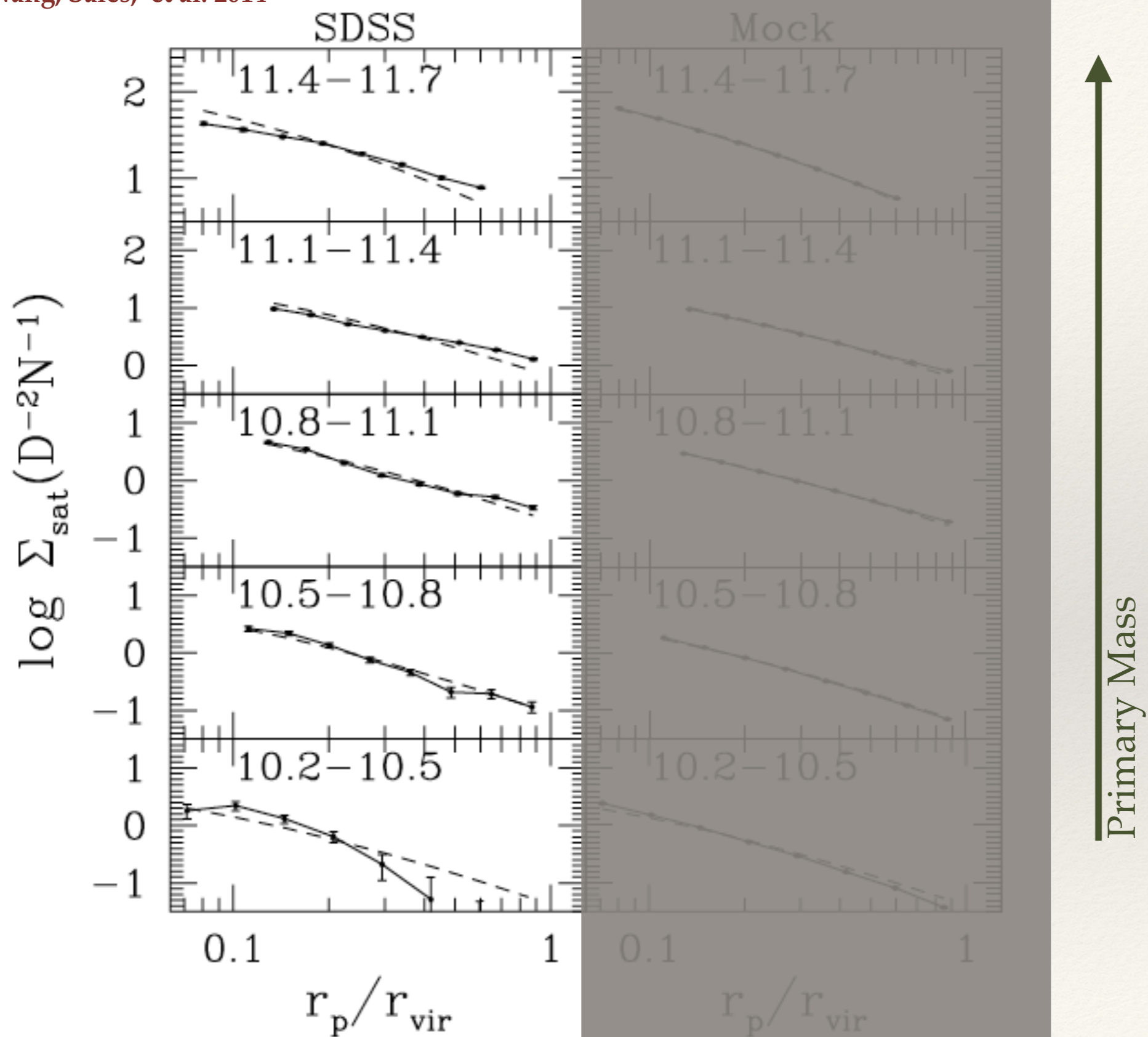
1. Select isolated primaries from spectroscopic survey
2. Look for all objects in photometric SDSS within estimated r_{vir}
3. Remove the average background counts
4. Left over signal characterizes the satellite population (excess count, profile, color distribution, etc)



Repeat for Millennium I&II + SAM

Satellite Radial Distribution around Isolated Primaries

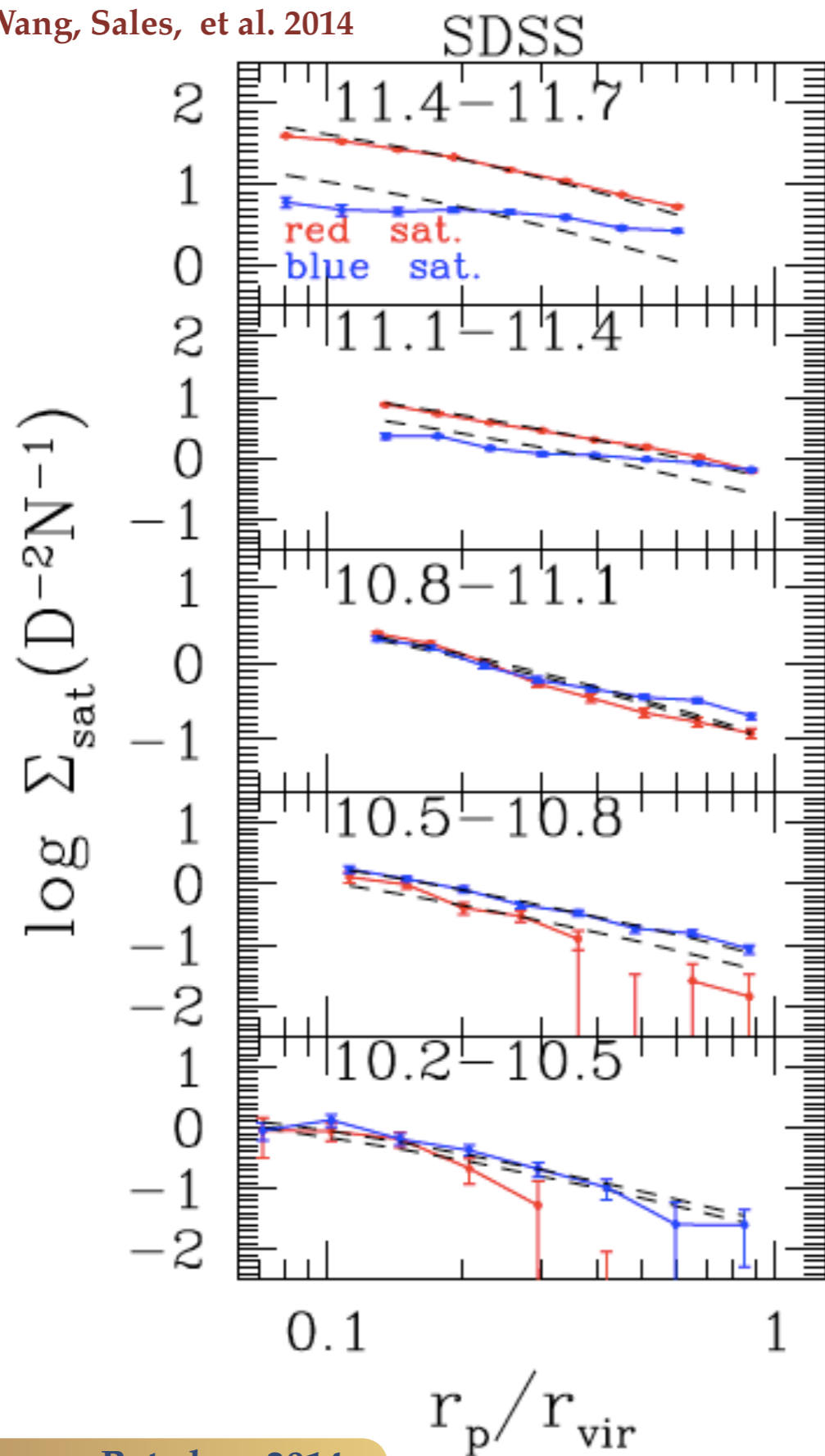
Wang, Sales, et al. 2014



Satellites roughly follow the underlying DM distribution

Satellite Radial Distribution according to color

Wang, Sales, et al. 2014



Annotations for the radial distribution plots:

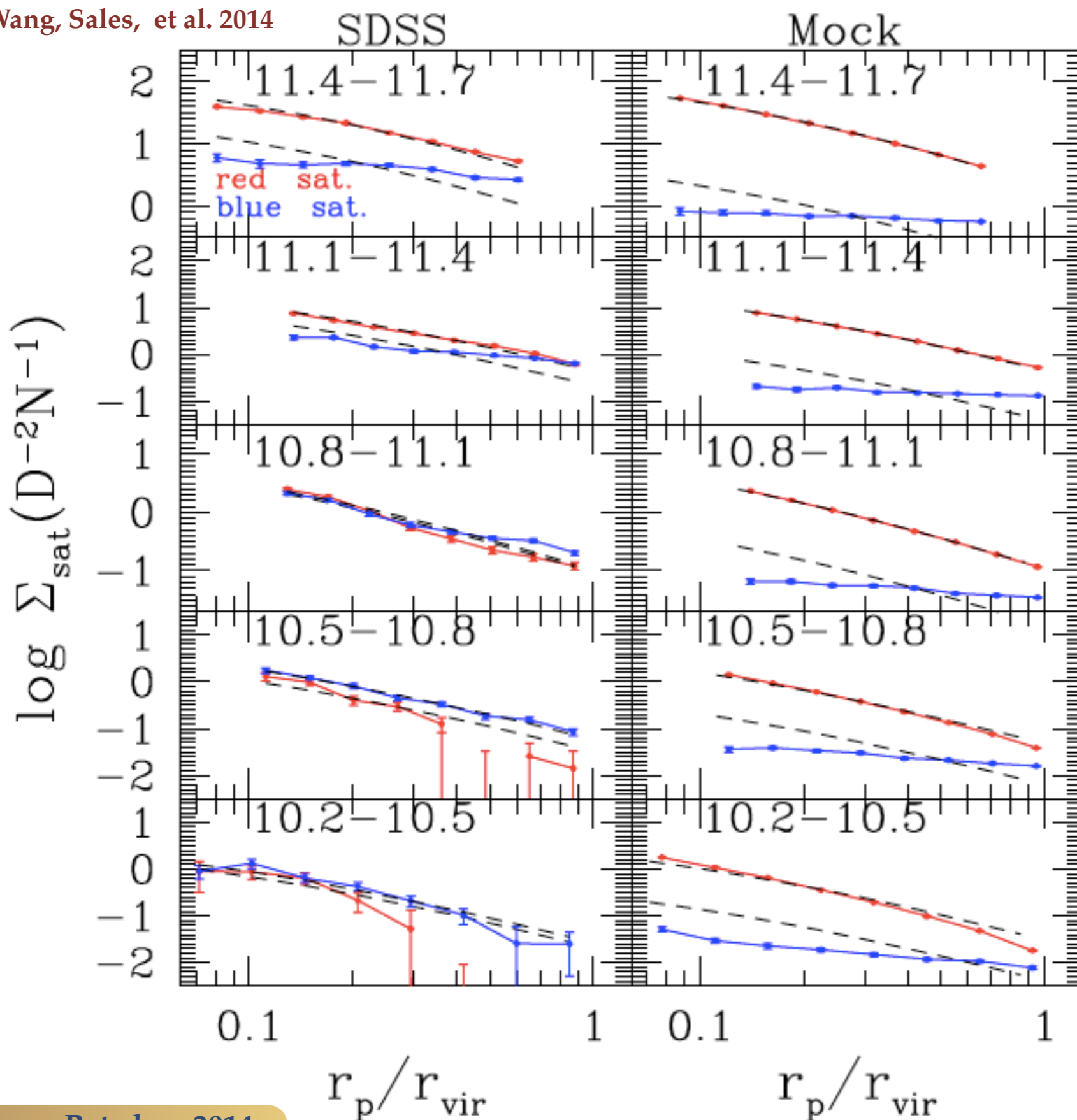
- Dominant red**: Points to the upper panels (brighter magnitude bins) where red satellites are more numerous.
- Flat profile for blue sats.**: Points to the blue satellite profiles in the upper panels, which are flatter than the red profiles.
- Dominant blue**: Points to the lower panels (fainter magnitude bins) where blue satellites are more numerous.
- Comparable slopes for red & blue**: Points to the profiles in the lower panels where the radial distributions of red and blue satellites are similar.

Primary Mass

Satellite Radial Distribution according to color

Wang, Sales, et al. 2014

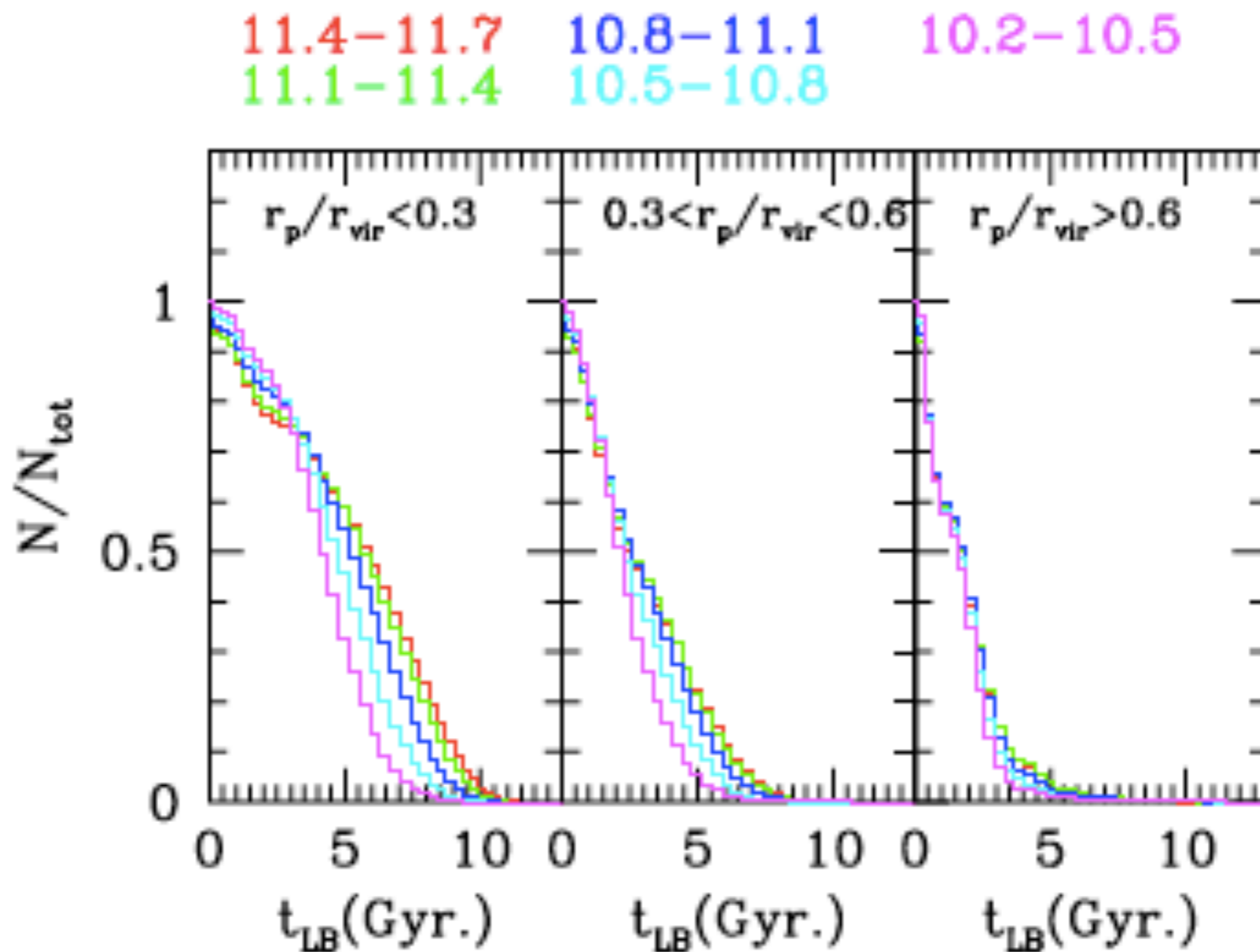
Primary Mass



SAMs overproduce red satellites
&
blue population has a shallow
profile

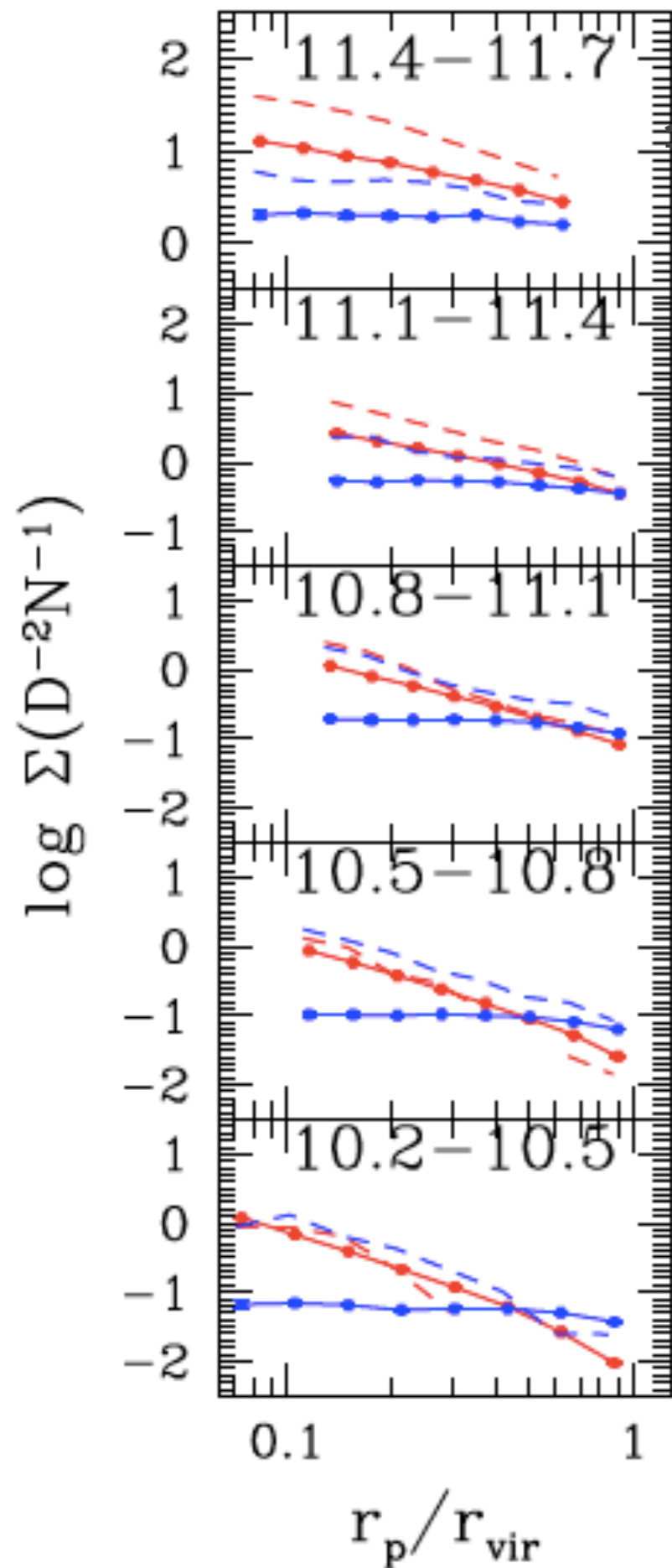
(see also: Weinmann 2006,2010; Kang 2008; Kimm 2009; Guo et al. 2013)

Satellite Infall Times



Wang, Sales, et al. 2014

Satellites in the inner regions fall in ~ 5 Gyr ago. They seem to remain blue in low mass primaries



Dashed = SDSS

Solid = SAM - Ram Pressure below $10^{14} M_{\text{sun}}$

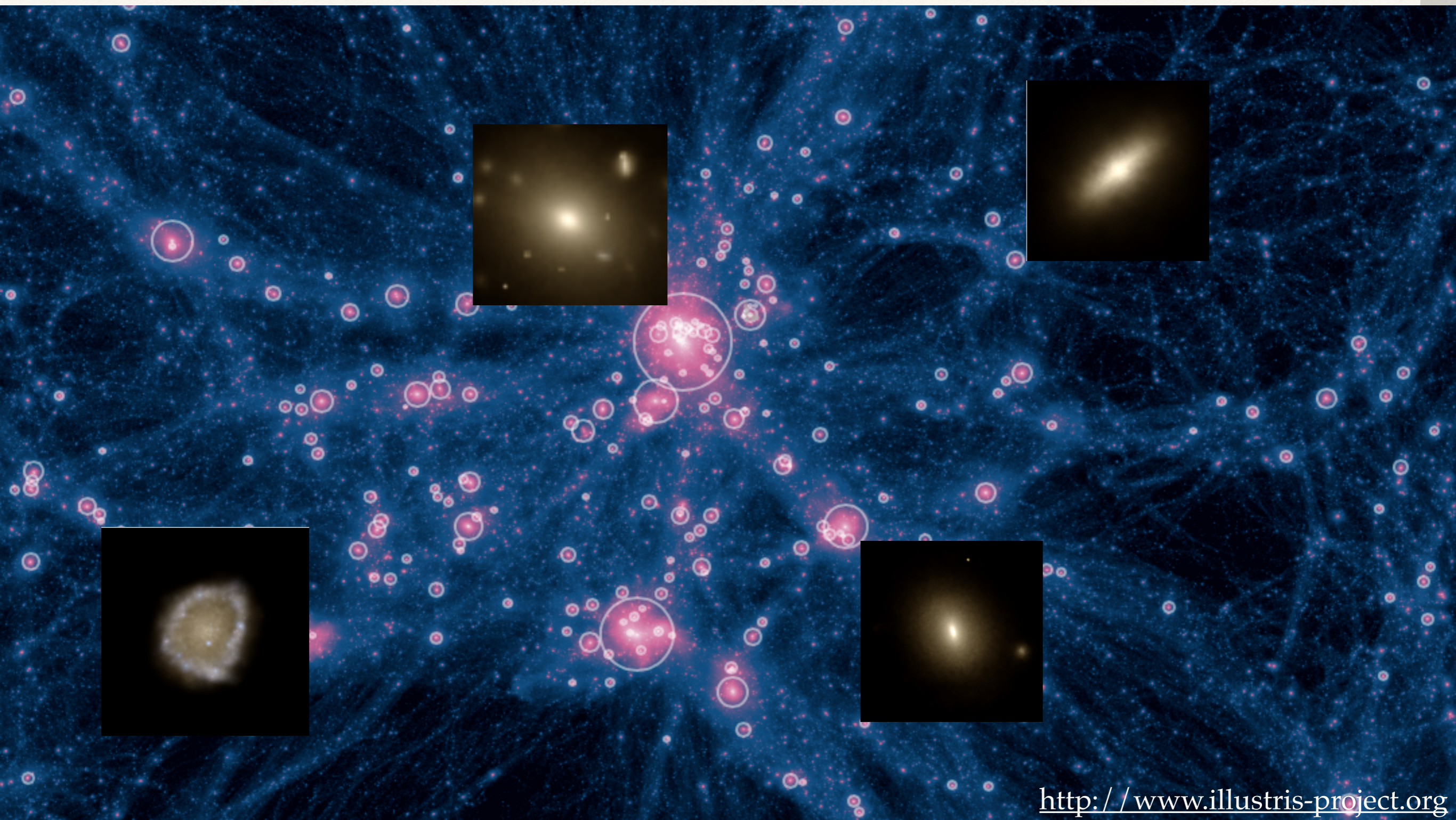
Artificially removing all environmental effects below $M_{\text{host}}=10^{14} M_{\text{sun}}$ does not solve the problem

(Increased fraction of blue satellites but radial profile remains too shallow)

Possible explanations:

- Spatial distribution of star formation?
- More complex re-accretion of gas?
- Infall properties?
- More?...

Satellite Galaxies in the Illustris Simulation

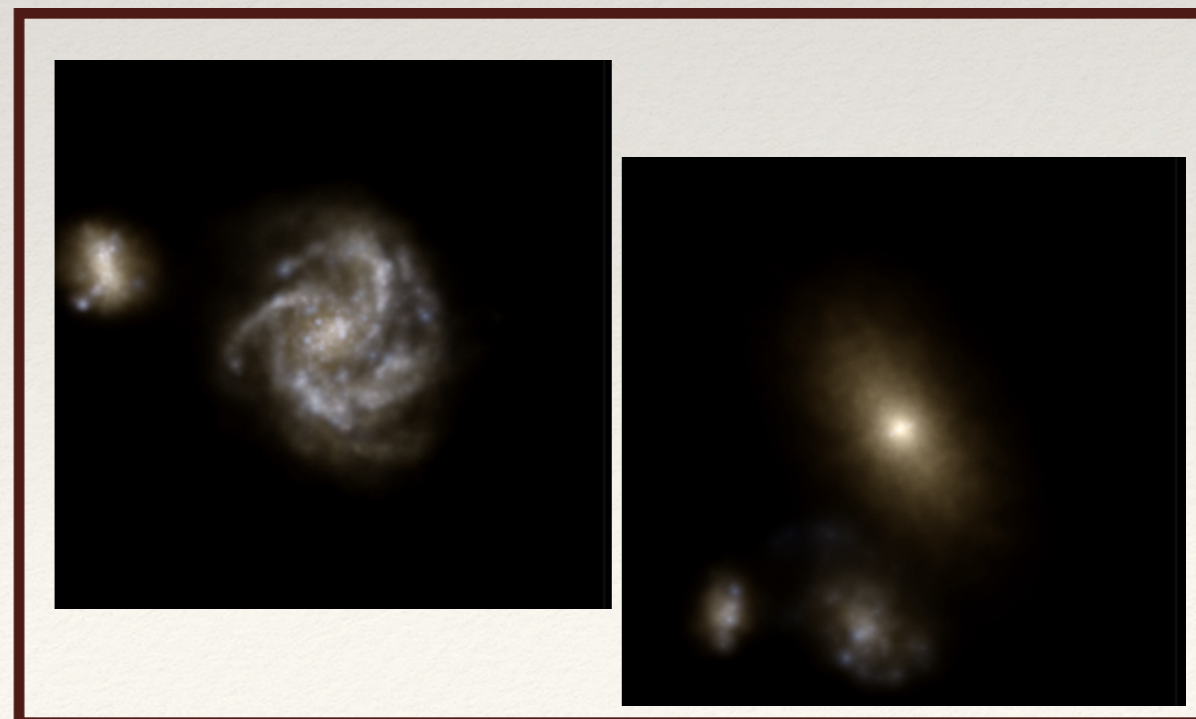


<http://www.illustris-project.org>

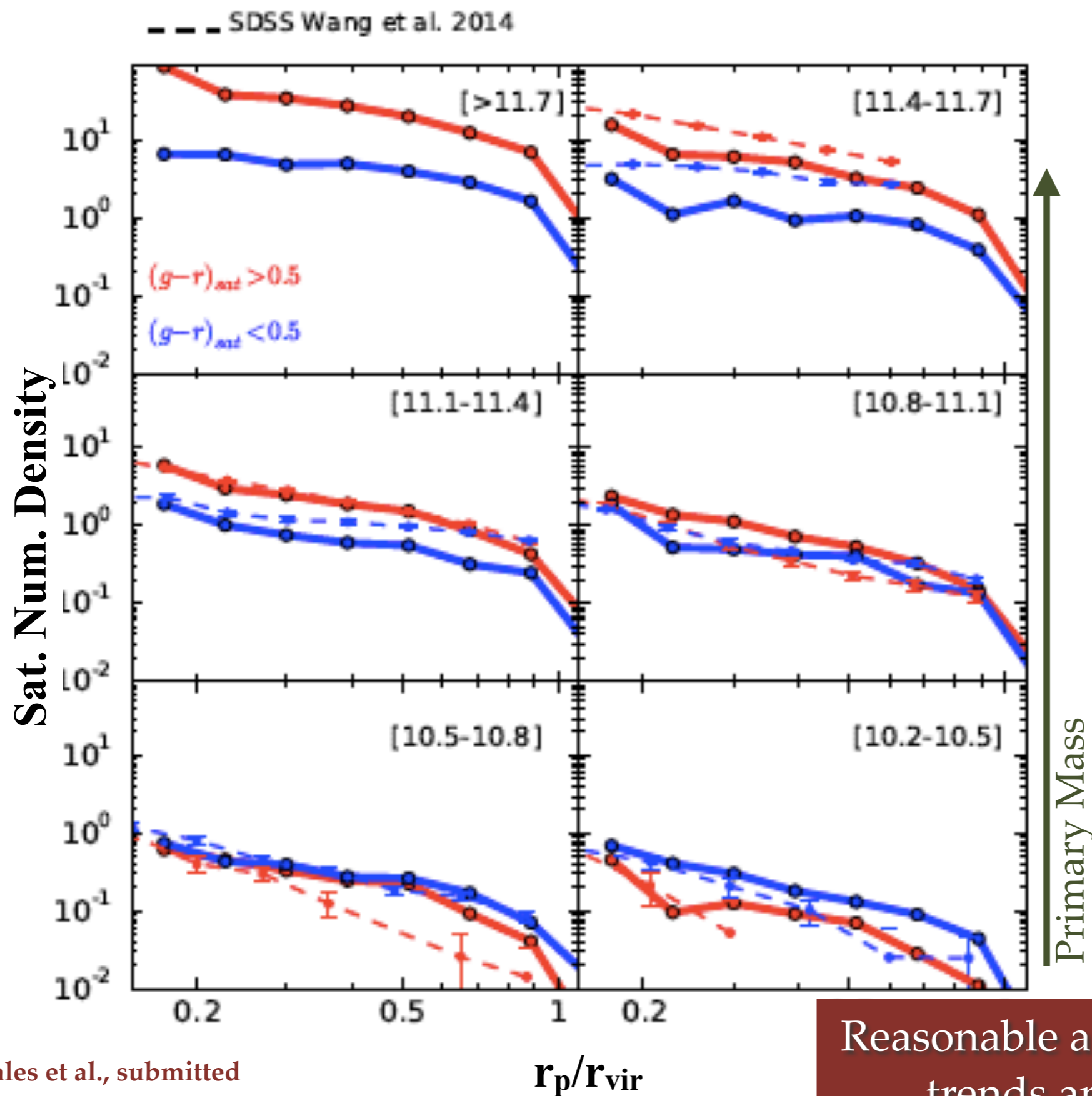


Satellites around massive primaries in Illustris

Satellites around MW-like primaries in Illustris



Satellite Radial Distribution according to color

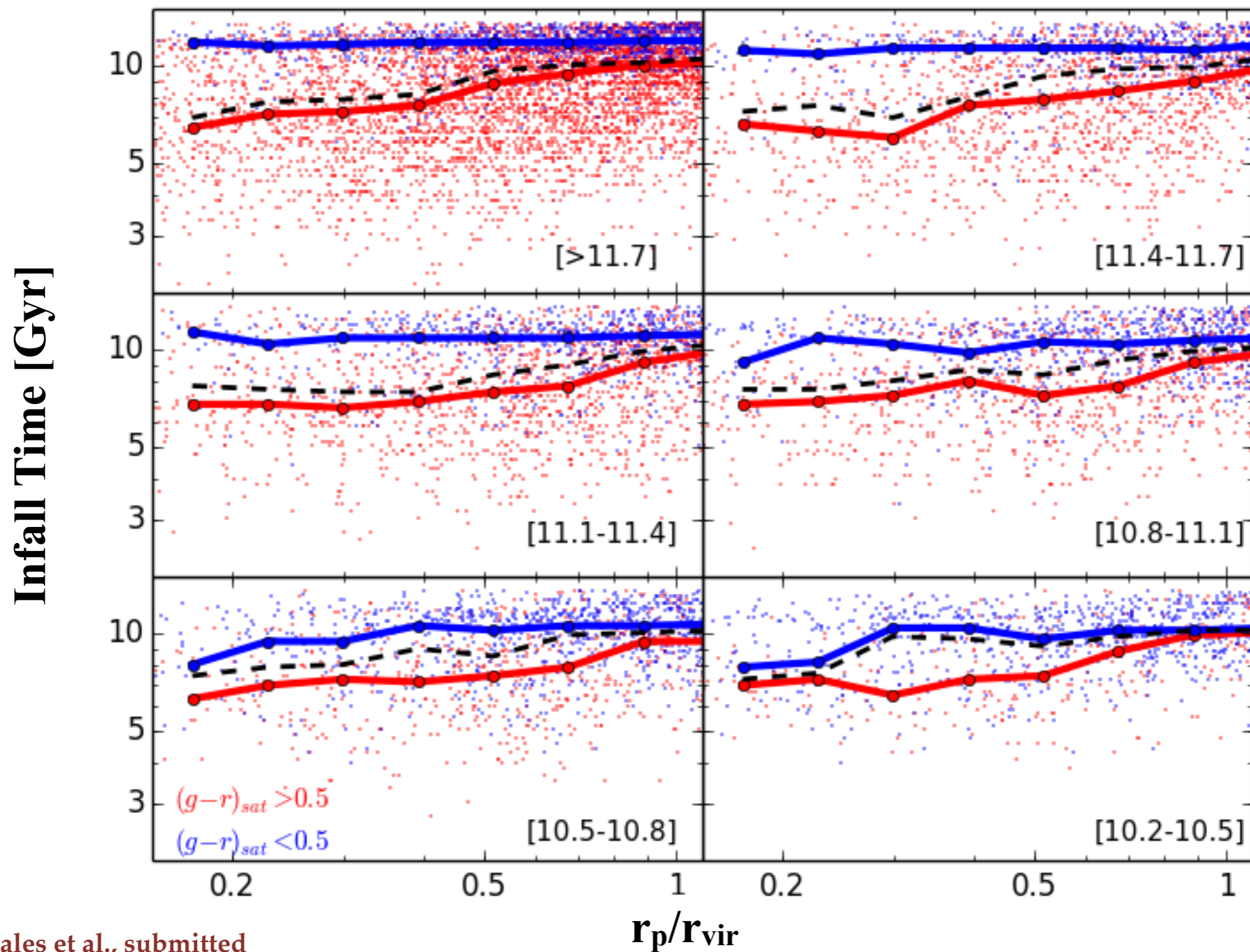


Sub-dominant & shallow blue distribution

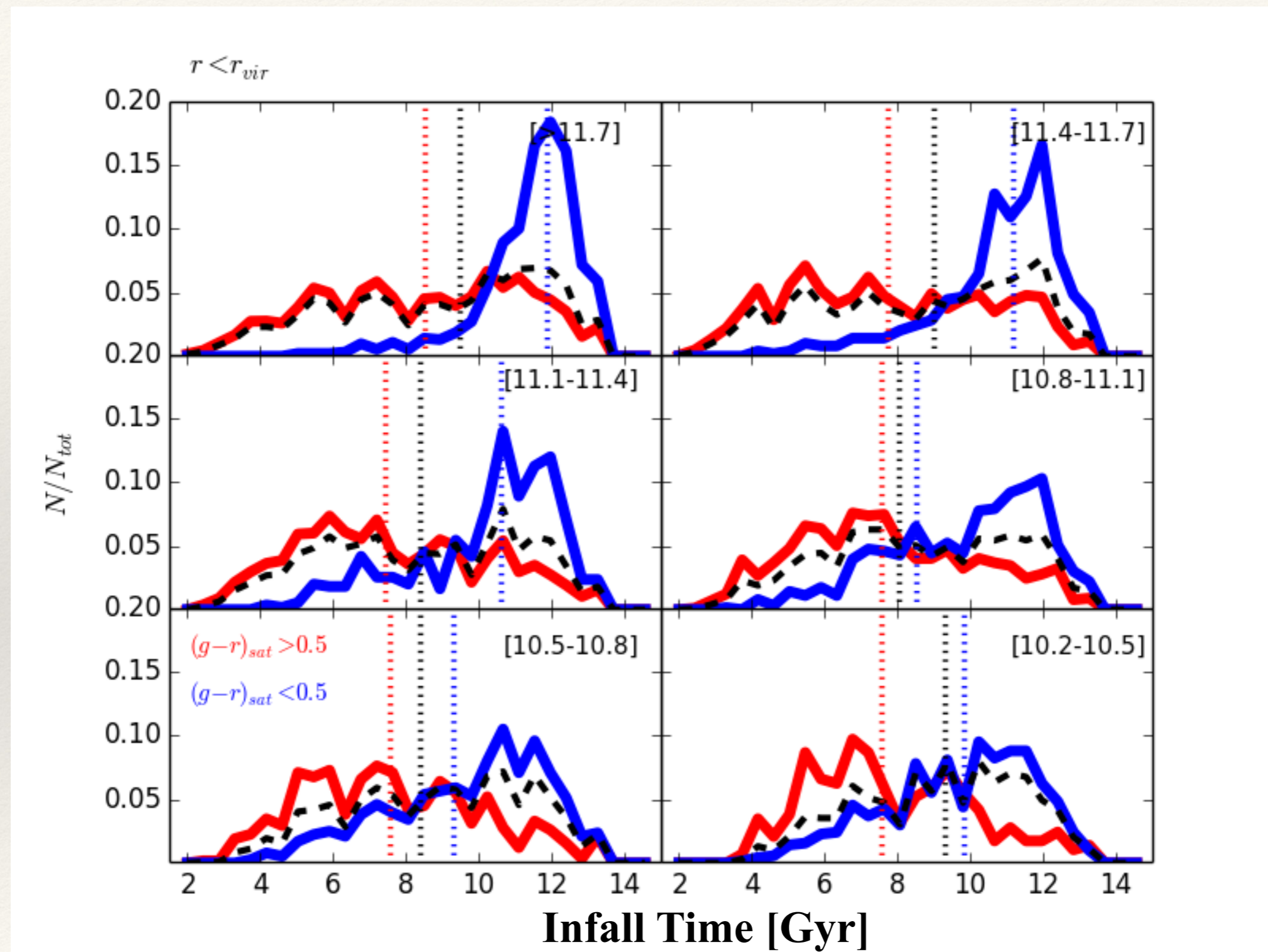
Dominant & steep blue distribution

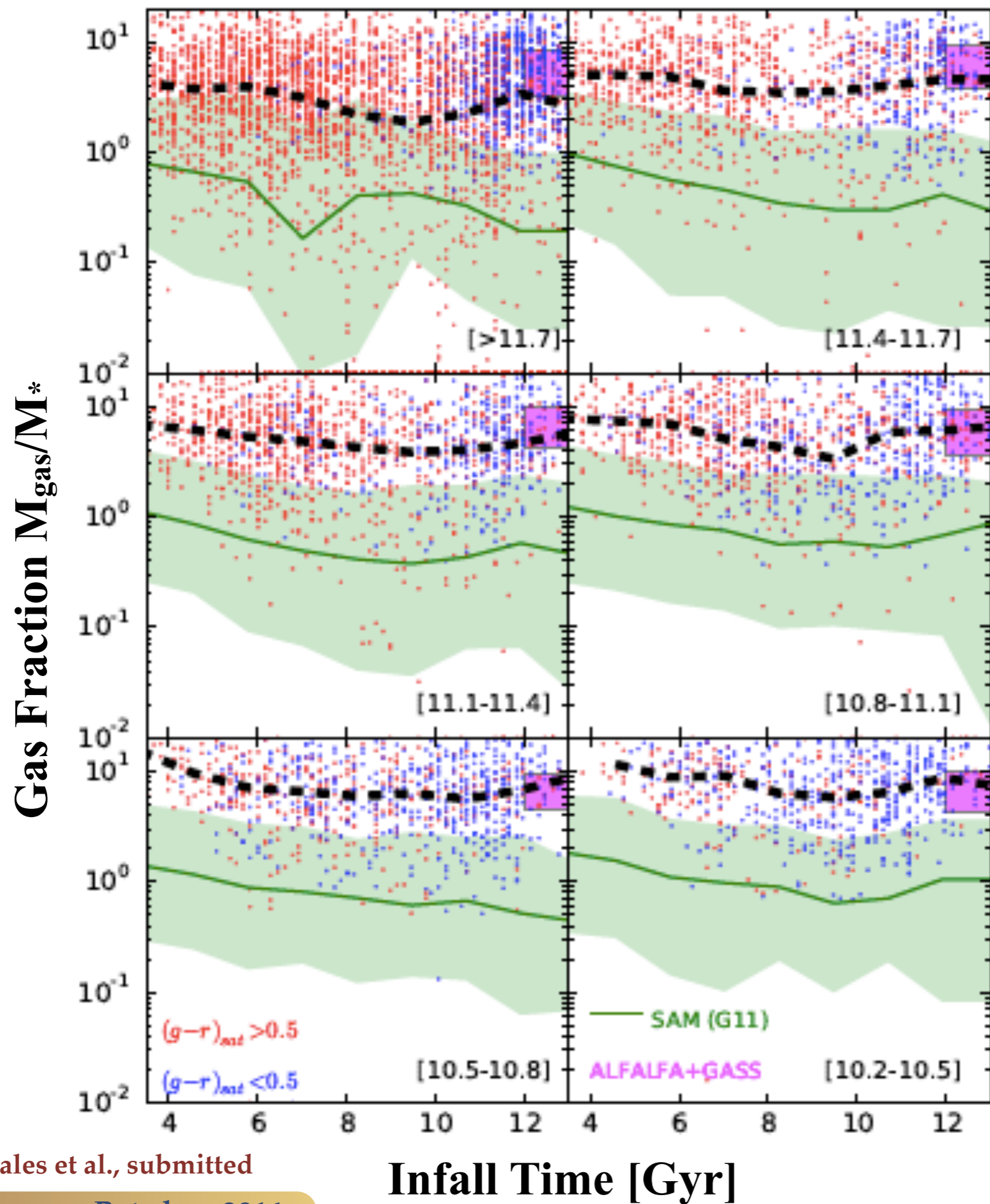
Reasonable agreement between observed trends and the hydro simulation

Satellite Infall Times & Colors in Illustris



Satellite Infall Times & Colors in Illustris

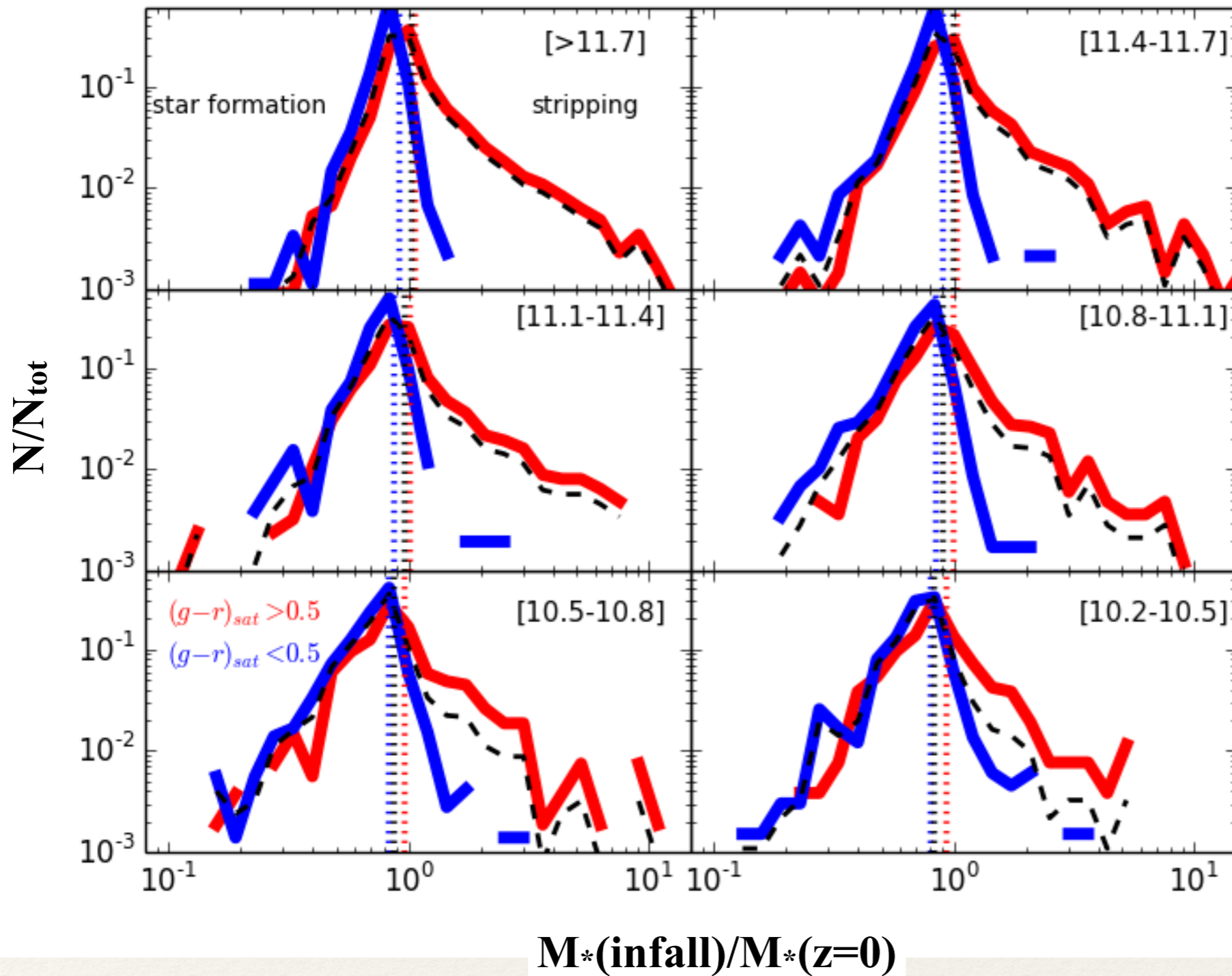




* Gas fractions of infalling dwarfs are significantly larger in Illustris than in previous SAMs

* Gas fractions vary weakly with redshift

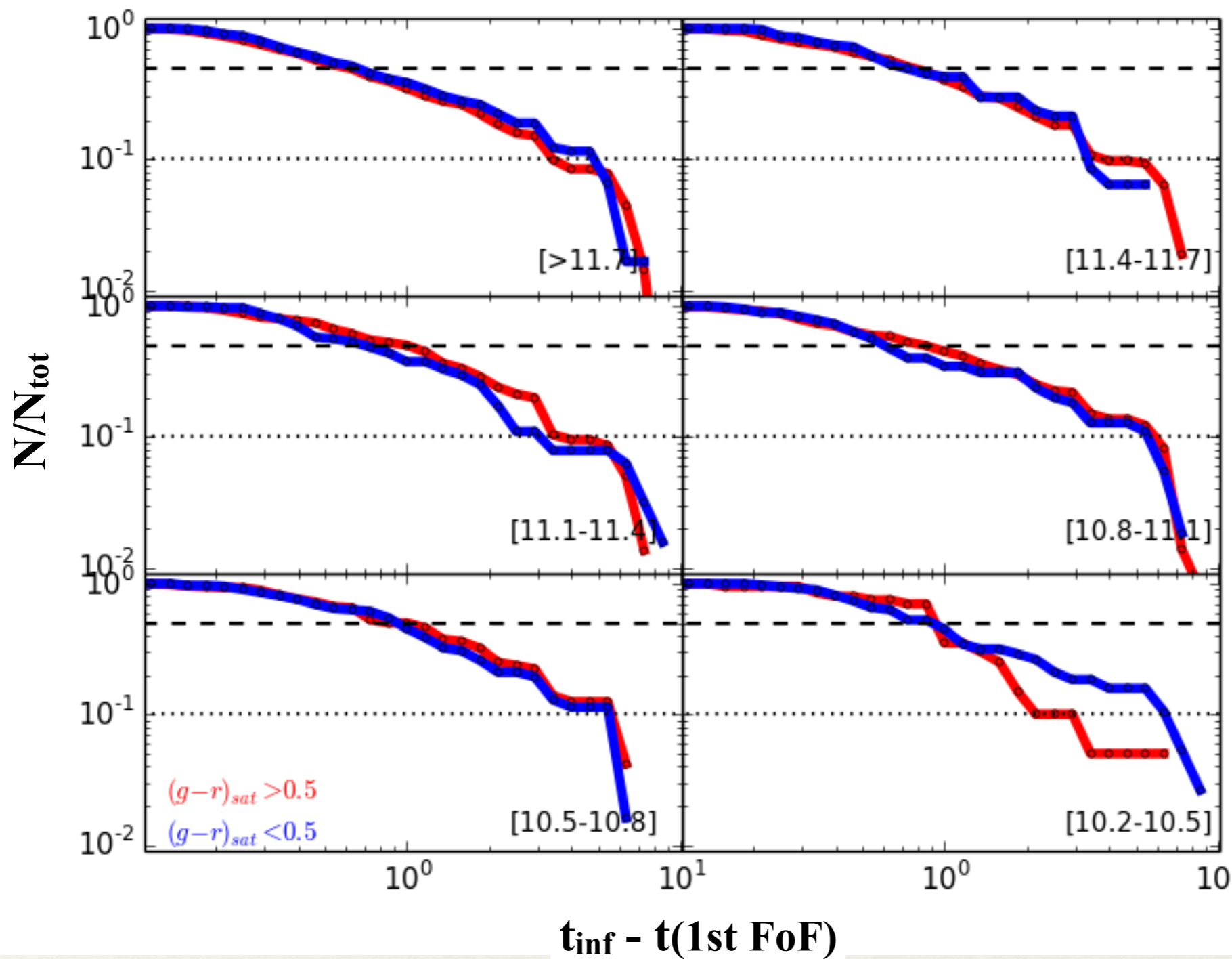
Satellite Stellar Mass Evolution



* On average, satellites have approx. same M^* at infall than at $z=0$

* 1%-10% of satellites more than double their mass while they are satellites

A word about pre-processing...



* ~50% satellites spent more than 1Gyr in a different group before joining the final host

* Pre-processing doesn't have a strong impact on the final satellite color

Summary

- ❖ In observations, **satellite distributions vary with primary mass**. Massive primaries have a dominant population of red satellites and blue objects distribute in a shallower profile. Instead, for MW-like objects, blue satellites are dominant and have a comparable slope to the red population
- ❖ Environmental effects weak-to-negligible for host masses $M_{\text{vir}} < 10^{13} M_{\text{sun}}$
- ❖ Reasonably good agreement in the color & distribution of satellites in the Illustris hydro-simulation
- ❖ **Large gas fractions** at infall seem key
- ❖ Strong suggestion of a **characteristic mass**: dwarfs with $M^* < 10^8 M_{\text{sun}}$ are all quenched in the Local Group but star forming in the field (see Wheeler et al. 2014, Geha et al. 2013)